# Erratum to the Paper «Improved Fits to the $x F_{3}$ CCFR Data at the Next-to-Next-to-Leading Order and Beyond» by A.L.Kataev, G.Parente, A. V.Sidorov (Physics of Elementary Particles and Atomic Nuclei. 2003. V.34. P.43-87) 

The bug crept into the calculations of the numerical values of the $A_{s}^{3}$ coefficients $C_{F_{3}}^{(3)}(n)$ in the QCD expression for the coefficient function $C_{F_{3}}^{(n)}\left(A_{s}\right)$ (the definition see in Eq. (8)) of the odd Mellin moments of Eq. (1) for $x F_{3}$ structure function of deep-inelastic neutrino-nucleon scattering $\left(A_{s}=\alpha_{s} /(4 / \pi)\right.$ ) with $n=1,3,5,7,9,11,13$ and $f=4$ numbers of flavours. This bug resulted from using the computer subroutine, which calculated the values for $C_{F_{3}}^{(3)}(n)$ from the given in Eq. (22) order $O\left(A_{s}^{3}\right)$ approximations for $C_{F_{3}}^{(n)}$, where instead of $f^{2}$ in the last terms $f$ was typed. These errors affected also the values of the even $C_{F_{3}}^{(3)}(n)$ coefficients, obtained from the explicitly calculated ones using the smooth interpolation procedure. The corrected results are given below in the 5th corrected column of Table 2 of the paper.

Table 2. The values for NLO , $\mathrm{NNLO}, \mathbf{N}^{3} \mathrm{LO}$ QCD contributions to the coefficient functions, used in our fits, and the results of $\mathbf{N}^{3} \mathrm{LO}$ Padé estimates

| $n$ | $C_{F_{3}}^{(1)}(n)$ | $C_{F_{3}}^{(2)}(n)$ | $\left.C_{F_{3}}^{(2)}(n)\right\|_{\text {int }}$ | $\left.C_{F_{3}}^{(3)}(n)\right\|_{\text {int }}$ | $\left.C_{F_{3}}^{(3)}(n)\right\|_{[1 / 1]}$ | $\left.C_{F_{3}}^{(3)}(n)\right\|_{[0 / 2]}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | -4 | -52 | -52 | -780.6427 | -676 | 480 |
| 2 | -1.778 | -47.472 | $(-46.4295)$ | $(-1206.83008)$ | -1267.643 | 174.4079 |
| 3 | 1.667 | -12.715 | -12.715 | -992.198975 | 97.00418 | -47.01328 |
| 4 | 4.867 | 37.117 | $(37.0076)$ | $(-269.865143)$ | 283.0851 | 246.0090 |
| 5 | 7.748 | 95.4086 | 95.4086 | 851.838501 | 1174.835 | 1013.328 |
| 6 | 10.351 | 158.2912 | $(158.4032)$ | $(2286.68115)$ | 2420.569 | 2167.903 |
| 7 | 12.722 | 223.8978 | 223.8978 | 3967.71313 | 3940.284 | 3637.790 |
| 8 | 14.900 | 290.8840 | $(290.8421)$ | $(5844.3042)$ | 5678.657 | 5360.371 |
| 9 | 16.915 | 358.5874 | 358.5874 | 7879.04004 | 7601.721 | 7291.305 |
| 10 | 18.791 | 426.4422 | $(426.5512)$ | $(10044.4785)$ | 9677.391 | 9391.308 |
| 11 | 20.544 | 494.1881 | 494.1881 | 12319.7676 | 11885.25 | 11633.28 |
| 12 | 22.201 | 561.5591 | $(561.2668)$ | $(14687.1133)$ | 14204.22 | 13991.80 |
| 13 | 23.762 | 628.4539 | 628.4539 | 171728.1191 | 16620.99 | 16449.68 |

The application of the corrected numbers in the next-to-next-to-leading order Jacobi polynomial fits of the experimental data of the CCFR collaboration resulted in slight decrease of $\mathrm{N}^{3} \mathrm{LO}$ values for $\Lambda \frac{(4)}{\mathrm{MS}}$ as presented in Tables 6,11 , and 12 by 3 MeV only and does not affect any conclusions of the paper.

